1 Application of Classical Kalman filtering technique in assimilation of

2 multiple data types to NeQuick model

- 3
- Patrick Mungufeni¹, Yenca Migoya-Orué², Tshimangadzo Merline Matamba³, George
 Omondi⁴
- 6
- ⁷ ¹Physics department, Muni University, P.O. Box 725 Arua, Uganda

²The Abdus Salam International Centre for Theoretical Physics (ICTP), Strada Costiera
 11, 34151, Trieste, Italy

- ³South African National Space Agency (SANSA), P.O. Box 32, Hermanus, 7200, South
 Africa
- ⁴Department of Physics and Materials Science, Maseno University, Kenya
- 13

14 Abstract

This study attempts to improve estimation of ionospheric electron density profiles over 15 Korea and adjacent areas by employing classical Kalman filtering technique to 16 assimilate Total Electron Content (TEC) data from various sources into the NeQuick 17 model. Successive corrections method was applied to spread the effect of TEC data 18 assimilation at a given location to others that lacked TEC observations. In order to 19 reveal that the assimilation results emulate the complex ionospheric changes during 20 aeomagnetic storms, the selected study days included both quiet (Kp \leq 3) and disturbed 21 22 geomagnetic conditions in the year 2015. The results showed that assimilation of TEC data derived from ground-based GPS receivers can improve the root mean squared 23 24 error (RMSE) associated with the NeQuick model estimation of ionospheric parameters by \geq 56 %. The improvement of RMSE achieved by assimilating TEC data that were 25 26 measured using ionosondes was ~50 %. The assimilation of TEC observations made by the COSMIC radio occultation technique yielded results that depicted RMSE 27 improvement of > 10 %. The assimilation of TEC data measured by GPS receiver 28

onboard Low Earth Orbiting satellites yielded results that revealed deterioration of
RMSE. This outcome might be due to either the fact that the receivers are on moving
platforms and these dynamics might have not been accounted for during TEC
computation or limitation of the assimilation process. Validation of our assimilation
results with global ionosphere TEC data maps as processed at the center for orbit
determination in Europe (CODE) revealed that both depicted similar TEC changes,
showing response to a geomagnetic storm.

Keywords: Ionosphere, modeling, data assimilation, NeQuick, Geomagnetic storms

10 **1. Introduction**

11

The peak electron density in the F2-region (NmF2) and total electron content (TEC) are 12 widely used parameters to characterize the ionosphere (Rishbeth and Gariott, 1969; 13 14 Gerzen et al. 2013). The NmF2 affects high-frequency (3 – 30 MHz) radio wave communication applications. For instance, Geeta and Yudav, (2014) stated that for 15 16 frequencies lower than 30 MHz, the ionosphere acts as a helpful aid for the radio wave propagation, but for the frequencies slightly greater than 30 MHz, it may cause 17 18 attenuation. The TEC affects Global Navigation Satellite System (GNSS) based positioning by introducing ionospheric refraction (Hofmann-Wellenhof et al., 2007) in 19 20 which the code delay or carrier phase advance occurs, resulting into a pseudo-range measurement instead of a true range measurement. 21

22 Several research efforts have tried to model NmF2 and TEC. For instance, the 23 Committee on Space Research and the International Union of Radio Science formed a 24 working group in the late sixties to produce an empirical standard model of the 25 ionosphere, the International Reference Ionosphere (IRI), based on all available data 26 sources (Bilitza et al. 1993). Output of the IRI includes NmF2 and TEC, among other 27 parameters. The NeQuick is another global ionospheric model that can be used to 28 obtain NmF2 and TEC. The NeQuick model and its subsequent modifications (NeQuick

G and NeQuick 2) are a three-dimensional, time dependent ionospheric electron density
model developed by the Abdus Salam International Center for Theoretical Physics
(ICTP) in Trieste, Italy and the Institute for Geophysics, Astrophysics and Meteorology
of the University of Graz, Austria (Nava et al. 2008, and references therein).

In order to produce a good empirical model of the ionosphere over a region, there is need to have extensive observations to be used for constructing the model. It may be difficult to obtain adequate observations that can be used in modeling since maintaining stable and reliable ground-based instruments all over regions may be expensive or not practicable, particularly over the seas and desserts. It is now known that estimations of ionospheric parameters by empirical and theoretical models may be improved by assimilating observations to the model.

Bust et al., (2004) presented the lonospheric Data Assimilation Three-Dimensional 12 (IDA3D) algorithm which uses a three-dimensional variational data assimilation 13 technique (3DVAR). The IDA3D is capable of incorporating most electron density 14 related measurements including GNSS-TEC measurements, low-Earth-orbiting beacon 15 TEC, and electron density measurements from radars and satellites. Bust and Datta-16 Barua, (2013) stated that lonospheric Data Assimilation Four-Dimensional (IDA4D) is an 17 ionospheric data assimilation algorithm which provides global 3-D time-evolving maps of 18 ionospheric electron density. A computationally practical data assimilation technique 19 known as Best Linear Unbiased Estimator (BLUE) has been implemented by Angling 20 and Cannon (2004) for combining RO data with background ionospheric models. 21 Recently, Mengist et al. (2019), investigated the IDA4D technique over Korea and the 22 23 neighboring areas, considering IRI model as the background. They showed that assimilation of ground-based global positioning system (GPS) Slant TEC (STEC), 24 NmF2 obtained from Constellation Observing System for Meteorology, lonosphere and 25 Climate (COSMIC) radio occultation (RO) and ionosondes yielded the best results. 26 27 Ssessanga et al. (2019) presented a preliminary study that assessed the capability of their developed four-dimensional (in space and time) data assimilation scheme to more 28 29 accurately estimate the 3-D picture of the ionosphere over the South African region.

Yue et al. (2007) assimilated electron densities observed by the Millstone Hill incoherent
scatter radar (ISR) into a one-dimensional midlatitude ionospheric theoretical model by
using an ensemble Kalman filter (EnKF) technique.

The current study performed assimilation of TEC derived from ground-based GPS 4 reciever and ionosonde stations as well as COSMIC RO to the NeQuick model. 5 Moreover, the study assimilated the TEC derived from GPS receivers onboard Low 6 Earth Orbiting (LEO) Swarm and COSMIC satellites. The classical Kalman filtering 7 technique applied to assimilate TEC observations to NeQuick model and the successive 8 corrections method used to correct the NeQuick model generated ionospheric 9 parameters at locations and epochs that lacked TEC observations make this study 10 11 unique. After describing in section 2 the data used in this work, we present in sections 3 and 4 the classical Kalman filtering technique and the successive corrections method, 12 respectively. The results and discussions are presented in section 5, while the 13 conclusions are presented in section 6. 14

15

16 2. Data sources

During geomagnetic storms, the variations in zonal electric fields and composition of the 17 18 neutral atmosphere contribute significantly to the occurrence of negative and positive ionospheric storm effects (Risbeth and Garriot, 1969; Buonsanto, 1999). In order to 19 20 ascertain the ability of the assimilation results to emulate the complex ionospheric changes during geomagnetic storms, the assimilations were done during March 15 -21 22 20, June 4, 21 – 26, October 6 – 11, 27, and December 18 – 24, 2015. In these four different months, representing seasons of March equinox, June solstice, September 23 equinox, and December solstice, the recorded minimum Dst were -222, -204, -124, and 24 -155 nT, respectively. For purposes of comparing ionospheric parameters on guiet and 25 26 disturbed geomagnetic conditions, there was one quiet day (Kp \leq 3) in each of the selected months. The selected geomagnetically quiet (Kp \leq 3) days were March 15, 27 June 4, October 27, and December 18. Figure 1 presents the variation of Dst (panel 28 (a)'s) and Kp indices (panel (b)'s) during the period under study. The hourly Dst and 3 29

hourly Kp indices can be obtained from the World Data Center of Kyoto, Japan
(http://swdcwww.kugi.kyoto-u.ac.jp/).

The pattern of the variation of Dst shown in Figure 1 indicates several main and recovery phases of geomagnetic storms during the period mentioned. The high Kp index values (Kp > 5) on March 17 – 18, June 22 – 23, October 7 – 8, and December 20 - 21 confirm the occurrence of geomagnetic disturbances during the period. Therefore, the ionosphere might vary greatly in the period under study.



9 Figure 1: Variation of (a) Dst and (b) Kp during (1) March 15 – 20, 2015, (2) June 4, 21
 10 – 26, 2015, (3) October 6 - 11, 27, 2015, and (4) December 18 – 24, 2015.

11

Most global climatological ionospheric models such as NeQuick and IRI might not emulate possible rapid variations of the ionosphere due to geomagnetic storms. In order to improve estimation of electron density up to altitude of GPS satellites (~20200 km) during the previously stated disturbed geomagnetic periods, we performed data assimilation to NeQuick model driven by daily F10.7 values. It is important to mention
that unlike NeQuick model which can estimate electron density up to altitude of GPS
satellites, IRI only estimates it up to an altitude of 2000 km. Moreover, since NeQuick is
a quick-run model, it is suitable for data assimilation process as computation time might
be reduced.

Figure 2 shows with blue crosses and red diamonds the locations of the ground-based 6 GPS receivers and ionosonde stations, respectively, that were used to obtain the TEC 7 data. The stations indicated in the figure had available data during most of the days 8 considered in the study period. Table 1 provides the geophysical parameters (e.g., 9 geographic and geomagnetic coordinates) associated with the stations in Figure 2. The 10 11 Receiver Independent (RINEX) data format files of GPS receivers can be obtained from University NAVSTAR Consortium network (ftp://data-out.unavco.org). The RINEX files 12 were processed using a software described in Ciraolo et al. (2007), yielding TEC 13 together with other parameters such as time, elevation and azimuth angles, and 14 geographic longitude and latitude of the ionospheric pierce points. The software can 15 produce TEC data at resolutions of 30 seconds, 1, 5, 10 and 15 minutes. To reduce 16 computational burden of working with high resolution TEC data, this study considered 5 17 minutes' resolution data. 18

The TEC derived by integrating electron density profiles obtained from ionosonde 19 stations listed in Table 1 can be accessed from the National Oceanic and Atmospheric 20 Administration (NOAA) website via the link, ftp://ftp.ngdc.noaa.gov. The data obtained 21 from the NOAA website was in the form of auto-scaled ionospheric parameters such as 22 23 peak height in F2-region, foF2, and TEC which are stored in Standard Archiving Output (SAO) format files. The TEC data provided in SAO files have a resolution of 15 minutes 24 and are obtained by integrating electron density profiles up to altitude of ~700 km. 25 Reinisch and Huang, (2001) stated that the auto-scaling program (real-time ionogram 26 27 scaler with true height (ARTIST)) approximates the electron density profile above the F2 layer peak by an α -Chapman function with a constant scale height that is derived from 28 29 the bottom-side profile shape near the F2 peak. The ionospheric parameters in SAO

files are associated with flags showing confidence level which ranges from low (11) to
 high (55). The foF2 and TEC data obtained from the SAO files used in this study were

3 those with high confidence level (≤ 22).



- 5 Figure 2: Locations of GPS (blue crosses) and ionosonde (red diamonds) stations used in the
- 6 study
- 7 Table 1: Geophysical parameters of GPS and ionosonde stations used in the study

Station name	ID	Country	Geog lat (°)	Geog lon (°)	mag lat (°)				
GPS receiver stations									
Daejeon	DAEJ	S. Korea	36.40	127.37	30.62				
Koganei	KGNI	Japan	35.71	139.49	29.46				
Aira	AIRA	Japan	31.82	130.59	25.92				
Changchu	CHAN	China	43.79	125.44	38.15				
Shimosato Hydrographic	SMST	Japan	33.58	135.94	27.46				

Suwon	SUWN	S. Korea	37.27	127.05	31.51			
Hsinchu	TCMS	Taiwan	24.79	120.99	19.29			
Hsinchu	TNML	Taiwan	24.79	120.99	19.29			
Usuda	USUD	Japan	36.13	138.36	29.91			
Yongsan	YONS	S. Korea	37.54	127.00	31.78			
BJNM, NIM	BJNM	China	40.24	116.22	34.94			
Beijing Fangshan	BJFS	China	39.61	115.89	34.31			
Yuzhno-Sakhalinsk	YSSK	Russia	47.03	142.71	40.71			
Tsukuba 2-A	TSK2	Japan	36.11	140.09	29.85			
Chichijima-A	CCJ2	Japan	27.07	142.20	20.96			
Ionosonde stations								
Jeju	JJ433	S. Korea	33.43	126.30	27.70			
Kokubunji	TO536	Japan	35.70	139.50	29.45			
Icheon	IC437	S. Korea	37.10	127.50	31.32			
Okinawa	OK426	Japan	26.33	127.80	20.62			

1

As mentioned in section 1, TEC measurements as obtained from the GPS receivers 2 onboard Swarm and COSMIC satellites were also used. The Swarm constellation is 3 composed of three identical satellites, namely, Alpha (A), Bravo (B), and Charlie (C). 4 Detailed information about orbital characteristics of the Swarm satellites can be found in 5 Zakharenkova and Astafyeva, (2015). Each of the Swarm spacecraft carries a Precision 6 7 Orbit Determination (POD) antenna. The GPS signal phase measurements as obtained from this antenna can be used to estimate the line of sight TEC between Swarm and 8 GPS satellites. This line of sight TEC data can be freely downloaded from the European 9 Space Agency (ESA) website (http://www.earth.esa.int/swarm). Its resolution was 1 and 10 10 seconds after and before July 15, 2014, respectively. The COSMIC TEC data which 11 is based on POD antenna measurements is processed at 1 second resolution and 12

archived at the COSMIC Data Analysis and Archive Centre (CDAAC) (http://cosmicio.cosmic.ucar.edu/cdaac/index.html). The line of sight TEC (STEC) obtained from
Swarm and COSMIC satellites were converted to vertical TEC (VTEC) at the position of
the LEO satellites as in Zhong et al. (2015).

The TEC data resulting from integration of electron density profiles associated with 5 COSMIC RO used in this study were also obtained from CDAAC. The integrated 6 7 electron density (integration being done up to the altitudes of the COSMIC satellites) can be obtained from ionPrf files. The TEC associated with a particular electron density 8 profile was assigned to the geographic coordinate of NmF2 in the same file. The 9 electron density profiles are obtained by the Abel inversion of RO data, assuming local 10 11 spherical symmetry of the electron density in a large region (a few thousand kilometers in radius) around the ray path tangent points (Krankowski et al. 2011). This assumption 12 may not always be valid, and horizontal ionospheric gradients may significantly affect 13 the retrieved electron density profiles, in particular below the F-layer. In addition, the 14 15 geographical location of the ray path tangent points at the top and at the bottom of a profile may differ (horizontal smear) by several hundred kilometers. Several studies 16 (e.g. Krankowski et al., 2011 and Mengist et al., 2019) that have used COSMIC data 17 commonly consider measurements with horizontal smear > 1500 km prone to errors and 18 19 they reject such measurements. Typically, over an area bounded by 5 and 4 degrees' longitude and latitude ranges, respectively, the total number of COSMIC TEC data 20 obtained in a day may be ~3. Therefore, rejection of data may make this number reduce 21 further. The current study did not reject COSMIC TEC with horizontal smear > 1500 km 22 23 since Mungufeni et al. (2020) analyzed COSMIC TEC data which were coincident with TEC estimated by ionosonde stations over South Africa, finding that, compared to 24 measurements with horizontal smear > 1500 km, some measurements with horizontal 25 smear < 1500 km were far from the linear least squares fitting line. 26

27

28 **3 The classical Kalman filtering technique**

In this study, we considered TEC measurement y at time t_k to improve the NeQuick model estimate of the ionospheric electron density profile x along the path that contains electrons that constitute y. We further considered that y is linearly related to x via the equation (Grewal and Andrews, 2001; Angling and Cannon, 2004),

$$5 \quad y \quad = \quad Hx + w \tag{1}$$

6 where *H* is the measurement sensitivity matrix and *w* is the measurement noise. If *x* 7 consists of *p* electron density values, *H* will be a row matrix with dimension *p*. As 8 justified later in this section, elements of *H* were considered to be vertical grid spans 9 corresponding to electron density values that constitute *x*. The improved estimate of 10 electron density profile x_a based on assimilation of measurement *y* into the NeQuick 11 model profile x_b is given by (Angling and Cannon, 2004),

12
$$x_a = x_b + K(y - Hx_b)$$
 (2)

¹³ where *K* is the Kalman gain which can be determined as

14
$$K = BH^T (HBH^T + R)^{-1}$$
 (3)

In equation (3), *B* and *R* are the background and observation covariance matrices,
 respectively, while superscripts *T* and -1 denote transpose and inverse of the matrix,
 respectively.

In general, for *n* values of *y* within a period of 15 minutes and horizontal grid cell having
 geographic longitude and latitude spans of 5 and 4 degrees, respectively, equations 2
 and 3 can be written as

21
$$x_a^i = x_b^i + K^i (y^i - H_i x_b^i)$$
 (4)

23
$$K^{i} = B_{i}H_{i}^{T}(H_{i}B_{i}H_{i}^{T} + R_{i})^{-1}$$
 (5)

respectively, where $i = 1, 2, \ldots, n$.

As in Yue et al. (2007), measurement noise values *w_i* can be considered as white noise
with zero expectation so that *R_i* in equation 5 is

3
$$R_i = \operatorname{var}(w_i) = c_r y_i^2$$
, (6)

4 where $c_r = 0.01$ and y_i is the ith observation which is assimilated.

5 The background error was considered so that p by p matrix *B_i* consists of elements 6 determined as (Yue et al., 2007),

₇
$$B_i(u,v) = c_b x_b^i(u) x_b^i(v) e^{\frac{-H_i(u)}{L_i(u)}}$$
 (7)

8 where $c_b = 0.001$, u = 1, 2, 3, ..., p, and v = 1, 2, 3, ..., p. Following Bust et al. (2004), 9 the distances $L_i(u)$ for scaling down $H_i(u)$ were considered as 20 km in E- and F-regions

10 and 500 km in the plasmasphere. This implies that

11
$$\frac{H_{i}(u)}{L_{i}(u)} = \begin{pmatrix} \frac{H_{i}(u)}{20}, & \text{for height} \le 600 \ km \\ \frac{H_{i}(u)}{500}, & \text{for height} > 600 \ km \end{cases}$$
 (8)

Equation 4 corresponds to a specific grid cell and assimilation time window. This approach reduces the huge demand for computational resources especially when all the grid cells are considered at once as mentioned in Rodgers, (2000). This study termed the filtering as indicated in equation 4, as classical Kalman filtering due to its similarity to the application of the same in classical mechanics (e.g., prediction of river floods and tracking/navigation of ships and spacecrafts).

It is important to emphasize that y_i can be obtained from any of the 5 data sources discussed in section 2. The *y* values obtained from ground-based GPS receivers were treated similar to the TEC measurements of ionosonde and COSMIC RO which are considered vertical. This was done by limiting *y* values obtained from GPS receivers to slant TEC (STEC) observed at very high elevation angles (>60°). Moreover, the ionospheric pierce points associated with the STEC were restricted to that of the specific spatial grid cell considered. The first advantage of this procedure is minimization of multipath effects on TEC observations and the second is rendering simplicity to the assimilation process (*y* occupies one horizontal grid cell). In future we might adopt the method described in Angling and Cannon (2004) to treat low elevation angle TEC observations that cross several horizontal grid cells.

Since NeQuick model can yield electron density profile for heights starting from about 6 60 to 20,200 km (approximate GPS satellite altitude), the altitudinal intervals for 7 computing electron densities (x_b^i in equation 4) were varied in this range. These 8 altitudinal intervals which constitute elements of H_i were set based on the known typical 9 vertical electron density profile as follows; in E- and F-regions (<600 km), electron 10 11 densities were computed at intervals of 10 km, for altitude region of 600 – 2000 km, the interval was increased to 50 km and above 2000 km, the interval was further increased 12 to 2,000 km. Therefore, for the case of ground-based GPS receiver TEC, all the 13 elements of H_i were non zero. 14

In order to account for lack of consideration of the entire electron density profile within 15 the altitude range of 60 - 20,200 km during determination of y_i associated with TEC data 16 from Swarm and COSMIC satellites, COSMIC RO, and ionosondes, some elements of 17 H_i were set to zero. For instance, while considering measurements associated with 18 ionosonde and COSMIC RO, elements of H_i corresponding to heights >700 and >800 19 km, respectively were set to zero. For the case of Swarm A and C which fly at altitude 20 ~460 km, elements of H_i corresponding to heights below this altitude were set to zero. 21 While for Swarm B and COSMIC satellites, elements of H_i corresponding to heights 22 23 <510, and <800 km, respectively were set to zero.

It can be deduced from the sentence preceding equation 4 that TEC data assimilations
were done at horizontal grid cells which contained TEC observations. The influence of
TEC data assimilation at horizontal grid cells which lacked TEC data were achieved
through successive corrections method (Bergthorsson and Döös, 1955; Rodgers, 2000;
Bratseth, 1986).

1 4. The Successive Corrections Method

In order to correct NeQuick model generated electron density profile at horizontal grid cell *d* based on assimilation results x_a^f (*f* = 1, 2, 3, . . ., F) in the nearest neighborhood of *d*, the successive corrections method described in Bratseth, (1986) and Rodgers, (2000) was applied after modifying it as,

6
$$x_a^d = x_b^d + \left(\frac{1}{F}\right) \times \sum_f \times 10^{-r} \times (x_a^f - x_b^f)$$
, (9)

where x_a^d is the corrected electron density profile, x_b^f is the background electron density 7 profile associated with x_a^f , and x_b^d is the background electron density profile at the 8 horizontal grid cell d. In equation 9, the expression $\left(\frac{1}{F}\right)$ in the second term was 9 introduced in order to obtain the average effect of assimilations at F grid cells in the 10 nearest neighborhood of grid d. Overall, simplification of the expression that constitutes 11 the second term in equation 9 yields small quantities. These small quantities either 12 increase or reduce elements (depending on the sign of $(x_a^f - x_b^f)$) of the NeQuick model 13 generated electron density profile, x_b^d at grid *d*. 14

Although the expression $(x_a^f - x_b^f)$ in equation 9 contains elements which are much 15 lower than typical NmF2 value (~10¹² electrons/m²), the elements might still have a 16 factor of 10 raised to a number < 12 as a power. Based on this idea, the term 10^{-r} in 17 equation 9 was introduced to allow the quantity needed to correct background electron 18 density (x_b^d) to reduce as *r* increases. In fact, the term 10^{-r} in equation 9 dictates that at 19 large values of r, assimilation results do not (minimally) influence background electron 20 density profile at grid d. On the other hand, when r is small, the assimilation results 21 influence the background electron density profile maximally. 22

Actually, *r* quantifies effect of temporal (d = f) or spatial separation (at a fixed epoch) between grid cells *d* and *f*. The consideration of *r* as spatial or temporal separation was

based on the idea that the local time difference between 2 locations separated by 15° 1 longitude is about 1 hour. After performing several trials, this study established that to 2 achieve smooth variations of electron densities spatially and temporarily, r values 3 should vary in steps of 0.25. Moreover, since elements in the expression $(x_a^f - x_b^f)$ are 4 expected to be small (\leq hundreds) compared to typical NmF2, the first r value was 5 6 assigned to 1. Equation 10 describes the variation of r as a function of longitudinal difference between cells d and f $(d_{lon}(f))$, latitudinal difference between cells d and f 7 $(d_{lat}(f))$, and time difference between epoch with observation data and that which lacked 8 9 observation data (dt(f)).

$$r = \begin{pmatrix} r_f (4 \times (i-1) < |d_{lat}(f)| \le 4 \times i) = 1 + (i-1) \times 0.25; & for |d_{lon}(f)| < 5 \text{ and } |dt(f)| < 15 \\ r_f (5 \times (i-1) < |d_{lon}(f)| \le 5 \times i) = 1 + (i-1) \times 0.25; & for |d_{lat}(f)| < 4 \text{ and } |dt(f)| < 15 \\ r_f (15 \times (i-1) < |dt(f)| \le 15 \times i) = 1 + (i-1) \times 0.25; & for |d_{lon}(f)| < 5 \text{ and } |d_{lat}(f)| < 4 \end{pmatrix}$$
(10)

where the two vertical bars represent magnitude and *i* is one of the positive integers 11 12 which was specified after knowing the spatial or temporal difference between grid cells d and f. For example, when all conditions set in the top row of the right hand side of 13 Equation 10 are satisfied and $|d_{lat}(f)| = 6$ degrees, *i* will take value of 2. 14 Then, $r = r_f(4 < |d_{lat}(f)| \le 8) = 1 + (2 - 1)^* 0.25 = 1.25$. Generally, at a particular assimilation 15 16 time window, all grid cells that lacked observation data were corrected in two ways. In the first case, $d_{lon}(f)$ was restricted to ≤ 5 degrees, while r varied with $d_{lat}(f)$ as in the top 17 expression in right hand side of equation 10. In the second case $d_{lat}(f)$ was restricted to 18 \leq 4 degrees, while r varied with d_{lon}(f) as in the middle expression in right hand side of 19 equation 10. Concerning corrections at a particular grid cell (d = f) for epochs that 20 lacked observation data, the r values varied with dt(f) as in the bottom expression in the 21 right hand side of equation 10. 22

23 5. Results and discussions

24 **5.1 An example of data assimilation process**

Figure 3 presents an example of assimilation result $(x_a^1, x_a^2, \dots, x_a^n)$ based on 1 equation 4 and TEC observed by ground-based GPS receivers on October 7, 2015. The 2 data assimilated were those at grid cell centered at longitude and latitude 127.5° E and 3 38° N, respectively as well as time interval 0:00 – 0:15 UT. In this assimilation window, 4 we considered only 21 TEC observations (n = 21) to be assimilated. For clarity, we only 5 present in Figure 3 x_a^1 , x_a^2 , x_a^3 , x_a^{20} , and x_a^{21} which are associated with the first three 6 and the last two observed TEC values. The red line in Figure 3 represents background 7 electron density $(x_b^1, x_b^2, x_b^3, x_b^{30}, x_b^{21})$ obtained from NeQuick model. While 8 implementing equation 4, we set $x_b^1 = x_b^2 = \dots$, $x_b^n = x_b^n$. This implies that the 21 9 10 observed TEC values were treated as scalars and assimilated in a amanner similar to the recursive approach as in Grewal and Andrews, (2001). 11



Figure 3: Electron density profile obtained from NeQuick model (red) and improved
 profiles (., x, +, o, *) associated with assimilation of some TEC values within spatial grid

cell centered at longitude and latitude 127.5° E and 38° N, respectively. The date and
time interval associated with the data plotted are indicated on top of the panel.

Based on the similarity with the recursive approach, we considered x_a^{21} associated with the last assimilated observed TEC value as the assimilation result for the specific assimilation window. We need to mention that the peak electron densities computed from x_a^{21} and x_b^{21} and were 6.0×10^{11} and 8.5×10^{11} electrons/m³, respectively. These peak electron densities can be converted to foF2 using the expression (Davies, 1990),

$$8 foF2 = \sqrt{\frac{NmF2}{1.24 \times 10^{10}}} (11)$$

9 yielding 6.96 and 7.41 MHz, respectively. These two values of foF2 can be compared 10 with 5.94 MHz obtained from the ionosonde at IC437 (Lon 127°, Lat 36°) during the date 11 and assimilation window indicated in Figure 3. The comparison reveals that the 12 difference between foF2 obtained by assimilation procedure and the observed is smaller 13 than that between foF2 obtained from NeQuick model and the observed. For 14 generalization purposes several test scenarios were performed for the entire period 15 under study.

16 **5.2 Assimilation Test Scenarios**

The test scenarios presented in this subsection were validated at horizontal grid cells 17 centered at geographic latitude (longitude) (i) 26° (122.5°), (ii) 34° (137.5°), and (iii) 38° 18 (127.5°), (iv) 33.4°(126.3°), (v) 26.3°(127.8°), and (vi) 39.6°(115.9°). The validation at 19 grid cells (i) and (ii) were done using TEC data obtained from GPS receiver at TCMS 20 and foF2 data obtained from ionosonde at TO536, respectively, while that at (iii) was 21 done using TEC data obtained from GPS receiver at YONS as well as foF2 data 22 23 obtained from ionosonde at IC437. Moreover, the validation at grids (iv) and (v) were done using foF2 data obtained from OK426 and JJ433, while the validation at grid (vi) 24 was done using TEC data obtained from BJFS. It should be noted that validation station 25 data were not used during assimilation. 26

1 5.2.1 Assimilation of Ground-based GPS receiver TEC

We assimilated ground-based GPS receiver derived TEC to the NeQuick model using equation 4. The foF2 derived from the assimilation results over IC437 during March 15 -20, October 6 – 11, 27, and December 18 - 24, 2015 are presented in Figure 4, panels (a), (b), and (c), respectively. The corresponding foF2 observed by the ionosonde at the station as well as foF2 obtained from NeQuick model are superimposed. The blue, red, and black colors in Figure 4 and later in Figures 5 and 7 represent parameters obtained from NeQuick model, assimilation, and observation, respectively.

Figures 5 (a) - (d) present TEC over YONS obtained from assimilation, NeQuick model, 9 and GPS receiver station during March 15 - 20, June 4, 21 - 26, October 6 – 11, 27, and 10 December 18 - 24, 2015, respectively. The increase and reduction in comparison to a 11 background value of observed ionospheric parameters during geomagnetic storms are 12 usually termed as positive and negative ionospheric storm effects, respectively 13 (Buonsanto, 1999). As mentioned in section 2, this study associated the ionospheric 14 parameters during March 15, June 4, October 27, and December 18, 2015 when Kp ≤ 3 15 with background values. It can be deduced from Figures 4 (c), 5 (b) and (d) that values 16 of ionospheric parameters increased during the main phases of geomagnetic 17 disturbances in June and December solstices when compared to those on quiet days in 18 these seasons. Moreover, Figures 4 (a), (b), 5 (a) and (c) clearly show that values of 19 ionospheric parameters reduced significantly during the main phases of geomagnetic 20 disturbances in March and September equinoxes when compared to those on quiet 21 days in these seasons. Since detailed discussions about the physical mechanisms 22 23 responsible for the generation of positive and negative ionospheric storm effects are out of the scope of the current study, interested readers may refer to the previous studies 24 (e.g., Buonsanto, 1999; Mendillo and Klobuchar, 2006) for such discussions. 25

Generally, Figures 4 and 5 show that the assimilation yielded ionospheric parameters which are in most cases closer to the observed parameter than to those generated from NeQuick model. Therefore, the assimilation process yields ionospheric parameters which depict the response of the ionosphere to geomagnetic disturbance. Another

general feature of Figures 4 and 5 is that assimilation results and the observed data 1 exhibited daily/diurnal variability, which is almost not present in the climatological 2 3 NeQuick model values. However, visual inspection of Figure 4 shows some cases (e.g. panel (c) on December 18 and 24) where NeQuick model results are closer to the 4 observed data compared to assimilation results. These isolated cases might result 5 because the data assimilated are obtained using GPS receiver while the observation 6 data presented in Figure 4 are obtained using ionosonde. The inherent discrepancy in 7 the instruments might manifest in the observed cases of assimilation results not 8 performing well. The overall statistical analysis presented in Figure 6 (b) confirms that 9 these are isolated cases. 10



Figure 4: Panels (a), (b), and (c) present variation of foF2 during March 15 - 20, October
6 - 11, 27, and December 18 - 24, 2015, respectively over IC437. The blue, red, and

black colors represent NeQuick model generated parameter, assimilation result, and
 observed parameter, respectively.

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Figure 5: Panels (a) - (d) present variation of TEC during March 15 - 20, June 4, 21 –
26, October 6 - 11, 27, and December 18 - 24, 2015, respectively over YONS. The blue,
red, and black colors represent NeQuick model generated parameter, assimilation
result, and observed parameter, respectively.

Figures 6 (a) – (c) present the scatter plots of NeQuick model generated ionospheric parameter (red color) and assimilation results (black color) as a function of coincident

1 observed parameter over TO536, IC437, and YONS, respectively. It should be noted that the data plotted in Figure 6 is the same as that in Figures 4 and 5. The correlation 2 3 coefficients r, and root mean squared error (RMSE) associated with the data plotted are indicated on the respective panels. It can be seen in Figure 6 that r value associated 4 with assimilation result parameter over a particular station is always higher than that of 5 NeQuick model. Moreover, the RMSE value associated with assimilation result 6 parameter over a particular station appear significantly improved (reduced) compared to 7 that associated with NeQuick model. In fact, the improvement percentage (IP) of the 8 RMSE values were determined as 9

$$10 IP = \frac{NQ_E - ASM_E}{NQ_E} \times 100\% (12)$$

where NQ_E and ASM_E represent RMSE values associated with NeQuick model and assimilation result, respectively. The *IP* over TO536, IC437, and YONS were 69, 56, and 81 % respectively.



Figure 6: Panels (a) and (b) present scatter plots of foF2 obtained from NeQuick model 3 (red color) and assimilation result (black color) as a function of coincident observed foF2 4 5 over TO536 and IC437, respectively. Panel (c) presents TEC obtained from NeQuick model and assimilation result as a function coincident observed TEC over YONS. The 6 7 observations are those that fall within the study period

The RMSE improvement percentage values (\geq 56 %) we obtained in this study are 8 higher than the 44 % of Mengist et al. (2019) when they investigated the performance of 9 Ionospheric Data Assimilation Four-Dimension technique over Korea and adjacent 10 areas, considering International Reference lonosphere model as the background model. 11 The study by Mengist et al. (2019) was done during both geomagnetic quiet and 12 13 disturbed days (March 15 – 18, 2015). The limitation of the vertical grids to altitude of ~1,336 km in Mengist et al. (2019) might contribute to the low RMSE improvement 14 percentage obtained in their study compared to that of the current study. The study on 15 imaging South African regional ionosphere using 4D-var technique by Ssessenga et al. 16

(2019) might not reasonably estimate TEC up to altitude of GPS satellites since it
 limited the vertical grid to altitude of 1,336 km.

The results presented in Figures 4 and 6 signify that assimilation of TEC data obtained from ground-based GPS receiver to NeQuick model can be helpful in determining fairly well foF2 over a location that does not have ionosonde station.

6 5.2.2 Assimilation of TEC obtained from lonosonde stations

After assimilating ionospheric TEC data obtained from ionosonde stations to the 7 NeQuick model, the validation of the assimilation exercise was done using TEC data 8 obtained from YONS. It should be noted that most records of ionosonde stations did not 9 have available TEC data with the exception of IC437 which belongs to the same spatial 10 grid cell (ii) as YONS. Figure 7 presents the variations of TEC obtained from 11 assimilation (red) over YONS during the period under study. Superimposed over the 12 figure are the corresponding TEC obtained from NeQuick model as well as observed 13 TEC data obtained from GPS receiver over YONS. Figure 7 clearly shows similar 14 observations that were deduced from Figures 4 and 5. Overall, the TEC obtained by 15 assimilation closely follows the observed TEC. 16

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Figure 7: Panels (a) – (c) present TEC obtained from NeQuick model (blue),
assimilation (red), and observed by ground-based GPS receiver (black) during March
15 - 20, October 6 - 11, 27, and December 18 - 24, 2015, respectively over YONS.

In order to quantify how well the assimilation has improved estimation of TEC over 5 YONS, we present in Figure 8 a scatter plot of TEC obtained from assimilation (black) 6 and NeQuick model (red) as a function of coincident observed TEC over YONS. On the 7 figure, the r and RMSE associated with TEC obtained from assimilation and NeQuick 8 model are indicated. The r values portray that the TEC obtained from assimilation 9 correlates with observed TEC better than that obtained from background model. By 10 performing assimilation, the RMSE improved by ~50 %. This low improvement 11 compared to that associated with assimilation of ground based GPS receiver data could 12 13 be due to lack of inclusion of TEC above 700 km during assimilation. In addition, although electron density can be computed by NeQuick model up to an altitude of 700 14 km, the ionosonde observations may not in some cases reach this altitude. This might 15 partly explain why the assimilation TEC data in Figure 7 appears to be noisy. 16



Figure 8: Scatter plot of TEC obtained from NeQuick model (red color) and assimilation
result (black color) as a function of coincident observed TEC over GPS receiver station
at YONS. The data plotted are the same as that in Figure 7.

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It needs to be noted that the 50 % increment signifies that over locations which do not
have ground-based GPS receivers, assimilation of TEC obtained from ionosondes into
NeQuick model can be helpful in improving estimation of TEC data that would be
measured by ground-based GPS receivers.

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11 5.2.3 Assimilation of TEC obtained from COSMIC RO

Due to the scarcity of RO TEC data, we validated the results of assimilation of the data into NeQuick model with foF2 obtained from all the ionosonde stations considered in this study. Moreover, validations were done using TEC observed over Japan (KGNI),

South Korea (YONS), Taiwan (TCMS), and China (BJFS). Figure 9 (a) presents a 1 scatter plot of foF2 obtained from assimilation (black) and NeQuick model (red) as a 2 3 function of coincident observed foF2 over the ionosonde stations, while Figure 9 (b) presents TEC obtained from assimilation and NeQuick model as a function of coincident 4 observed TEC over the IGS stations. The r and RMSE associated with data plotted in 5 panels of Figure 9 are shown on the respective panels. Also indicated on the panels of 6 Figure 9 are the numbers of observed data which are coincident to assimilation results. 7 Despite considering 27 days in the year 2015 and all 4 ionosonde stations, there were 8 only 10 observations of foF2 values which were coincident with assimilation results. For 9 the case of validation with TEC data, there were only 17 observations which were 10 coincident with assimilation results. 11

Figure 9 exhibits that the assimilation of COSMIC RO TEC greatly improves estimation of foF2 and TEC since the r values associated with assimilation results are higher than those associated with NeQuick model. In fact, the RMSE improvement percentages for foF2 and TEC are 57 and 10 %, respectively. The percentage associated with TEC data is much less compared to that of foF2, maybe due to the limitation of COSMIC RO electron density profile altitude of ~800 km.



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Figure 9: Panel (a) presents scatter plot of foF2 obtained from NeQuick model (red) and
assimilation (black) as a function of coincident observed foF2 over the ionosonde
stations considered in this study. Panel (b) presents scatter plot of TEC obtained from
NeQuick model and assimilation as a function of coincident observed TEC over KGNI,
TCMS, YONS, and BJFS GPS receiver stations. The data plotted are those on the days
of the study period.

Based on the observations and discussions associated with Figure 9, the assimilation of
 COSMIC RO TEC seems to improve estimation of foF2 and TEC ionospheric
 parameters.

5.2.4 Assimilation of TEC obtained from GPS receivers onboard Swarm and COSMIC satellites

The results of assimilation of TEC obtained from GPS receivers onboard Swarm and COSMIC satellites were validated with ground-based GPS receivers over Japan (KGNI), South Korea (YONS), Taiwan (TCMS), and China (BJFS). Similar to Figure 9,

Figure 10 presents the scatter plot of NeQuick model TEC (red color) and assimilation 1 TEC (black color) as a function of coincident observed TEC over the ground-based GPS 2 receiver stations. Indicated on Figure 10 are the r and RMSE associated with the data 3 plotted in the figure as well as the number of coincident assimilation TEC and observed 4 TEC over the GPS receiver stations. Even though 27 days in the year 2015 and TEC 5 data from GPS receiver stations located in 4 different countries were considered, there 6 were still few (35) coincident data, as indicated in Figure 10. This observation may be 7 due to the fact that LEO satellites pass over a particular location reoccurs after several 8 days. 9

- 10 Figure 10 reveals that the assimilation of TEC obtained from GPS receivers onboard
- 11 LEO satellites yields lower correlation coefficient compared to that associated with the
- 12 NeQuick model. Actually, the RMSE deteriorated by ~107 %.



Figure 10: Scatter plot of TEC obtained from NeQuick model and assimilation as a
 function of coincident observed TEC over GPS receiver stations at KGNI, TCMS,
 YONS, and BJFS. The data plotted are those on the days of the study period.

The poor estimation of ionospheric parameters by the assimilation of TEC obtained from GPS receivers onboard LEO satellites need to be investigated further. However, we tentatively attribute this poor performance to the (i) dynamics of the receiver which might have not been considered in computing the POD TEC (ii) limitation of the assimilation technique.

In a practical application situation, assimilation of TEC data would depend on precedence of the data source which can be set based on the results depicted in the various test scenarios presented above. For instance, TEC data obtained from groundbased GPS receiver would be given the highest precedence, followed by TEC data from ionosonde, and lastly COSMIC RO TEC.

Based on the results from the test scenarios, for now we would not recommend 11 assimilation of TEC derived from GPS receivers onboard the LEO satellites to the 12 NeQuick model. If a horizontal grid cell does not have TEC data from any of the 3 13 recommended sources, the NeQuick model generated electron density profile at the cell 14 would be corrected as described in section 4. Examples of assimilation results obtained 15 from data assimilations followed by application of successive corrections method are 16 presented and compared with TEC data processed at the Center for Orbit 17 Determination in Europe (CODE) in section 5.3. As one of the international GPS service 18 (IGS) for geodynamics analysis centers, CODE provides daily Global lonospheric TEC 19 data Maps (GIMs) at www.aiub.unibe.ch/download/CODE/. 20

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22 **5.3 Validation of successive Corrections method**

This section presents in Figure 11 results where TEC data from all the ground-based GPS receiver and ionosonde stations indicated in Figure 2 as well as COSMIC RO were assimilated followed by successive corrections method to yield the final assimilation result. Later in this section we also present in Figure 12 similar results, but where TEC data from ionosonde stations at JJ433 and OK426 were not assimilated. These 2 stations appear to be suitable for validating further successive corrections method since

Figure 2 shows there were no ground-based GPS receivers within the grid cells that 1 contain the stations. In Figure 11, panels in columns (a) - (c) present CODE GIMs, TEC 2 obtained from assimilation result and NeQuick model, respectively. Panels in rows (i) -3 (v) present the TEC data at 5:00 UT (14:00 and 15:00 LT over Korea and Japan, 4 respectively) when considerable ionization is expected during October 6 - 10, 2015, 5 respectively. These dates were chosen to reveal the ionospheric changes before and 6 during the main phase (October 6 - 7, 2015) of the storm as well as during and after the 7 recovery phase (October 8 - 10). 8



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Figure 11: Panels in columns (a) – (c) present TEC obtained from CODE GIMs,
assimilation, and NeQuick model, respectively. Panels in rows (i) - (v) present the TEC
data at 05:00 UT during October 6 - 10, 2015, respectively.

It can be seen in Figure 11 that on October 7, 2015, TEC data from CODE and assimilation were the highest and reduced significantly on October 8, 2015. This observation is consistent with the results that were presented and discussed in section 5.2 (see Figures 4 - 6), where the TEC values during the main phase of the storm on

October 7, 2015 (see Figure 1) were found to be higher than that during the recovery 1 phase of the storm on October 8, 2015. As expected, these variations in TEC data 2 during the main and recovery phases of the storm are not strongly reflected in the 3 NeQuick model presented in Figure 11, panels in column (c). Visual inspection of 4 panels in row (iii) of Figure 11 shows that at longitude >132°, TEC data from CODE 5 GIMs are higher than that of assimilation and NeQuick. Since CODE GIMs are 6 constructed using a series of spherical harmonics functions whose coefficients are 7 determined using available TEC data from IGS stations (Schaer, 1999), CODE GIMs 8 TEC data over locations that lacked IGS stations may contain high error value. 9 Therefore, although TEC data from both CODE and assimilation seem to respond to 10 TEC changes due to occurrence of geomagnetic storm, TEC data from the two sources 11 may not perfectly correlate. Furthermore, although the resolution of the CODE GIMs 12 was changed from 2 hours to 1 hour on 19th October 2014, this hourly averaging might 13 still prevent the capturing of fine structures in the maps. 14

15 As mentioned before in this subsection, the successive corrections method was validated further using foF2 obtained from ionosonde stations at JJ433 and OK426. 16 Panels 1 (a) and 2 (a) in Figure 12 present the foF2 obtained from NeQuick model (blue 17 line), assimilation results (red line), and ionosonde stations (black dots). The 18 magnitudes of the differences between foF2 obtained from (i) NeQuick model and (ii) 19 assimilation results and the observed, denoted as Δ foF2 are presented in the panels 1 20 (b) and 2 (b) of the figure. It is important to mention that the data plotted in Figure 12 are 21 for the months and days in the month indicated on the horizontal axis of panel 2 (b). For 22 23 each day, the data were sampled at 4 hours interval. Particularly, the data corresponding to 02:00, 06:00, 10:00, 14:00, 18:00, and 22:00 LT are plotted in Figure 24 12. 25



Figure 12: Panels 1 (a) and 2 (a) show foF2 obtained from NeQuick model (blue line),
assimilation results (red line), and ionosonde stations (black dots) at JJ433 and OK426,
respectively. Panels 1 (b) and 2 (b) present the magnitudes of the differences between
foF2 obtained from NeQuick model (blue bars) and assimilation results (red bars) and
that observed at ionosonde stations at JJ433 and OK426, respectively.

It can be seen in panels 1 (b) and 2 (b) of Figure 12 that the Δ foF2 associated with 7 8 assimilation results are mostly smaller than those corresponding to NeQuick model. The 9 average AfoF2 associated with assimilation results over JJ433 and OK426 were established as 0.79 and 1.30 MHz, respectively, while the average Δ foF2 associated 10 with NeQuick results over JJ433 and OK426 were 1.04 and 1.40 MHz, respectively. 11 These results imply that over a particular station, the average Δ foF2 associated with 12 assimilation reduces significantly compared to that associated with NeQuick model. The 13 high average Δ foF2 and foF2 values over OK426 as depicted in Figure 12 might be 14 associated with the closeness of the station to the equatorial region where high 15 ionization and electro-dynamic processes occur. The sparse availability of ground-16

based GPS receivers within the vicinity of OK426 as shown in Figure 2 might be another reason for the high average Δ foF2 observed over the station. This is expected since the effectiveness of successive corrections method decreases as distance from locations of data increase.

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6 6. Conclusions

We made effort to improve estimation of ionospheric parameters over Korea and 7 adjacent areas by employing classical Kalman filtering technique to assimilate TEC data 8 from various sources into the NeQuick model. Successive corrections method was 9 applied to spread the effect of TEC data assimilation at a given location to others that 10 lacked TEC observations. The results from different assimilation scenarios showed that 11 data assimilation of ground-based GPS derived TEC data can improve root mean 12 squared error (RMSE) associated with the model estimation by ≥56 %. Assimilation of 13 14 TEC measured by ionosonde stations can improve RMSE associated with the model estimation of TEC data by ~50 %. The assimilation of TEC obtained from COSMIC RO 15 revealed RMSE improvement of ~10 %. Assimilation of TEC measured by GPS 16 receivers' onboard LEO satellites degraded the RMSE associated with the model 17 18 estimation by ~107 %, probably due to either the dynamics of the receivers or limitation of the assimilation technique. Validation of our assimilation results with global 19 ionosphere TEC data maps processed at CODE revealed that both reproduced similar 20 TEC changes, showing response to a geomagnetic storm. However, TEC data from the 21 22 two sources may not perfectly correlate.

For practical applications, we propose the assimilation of TEC data into the NeQuick model depending on the precedence of the data source which can be set based on the results presented in this study. That is, TEC data obtained from ground-based GPS receiver would be given the highest precedence, followed by TEC data from ionosonde, and lastly COSMIC RO TEC.

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